

Gauge Invariant Cosmological Perturbation Theory

Allison Pinto

PhD Advisor: Dr. Olaf Hohm

HU Berlin

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Motivation

Where do inhomogeneities in the universe come from?

How to construct a theory of gauge invariant cosmological perturbations?

What are the physical degrees of freedom that have an imprint in the CMB?

Outline

- 1 Electrodynamics
- 2 Gravity on Flat Space
- 3 Gravity on FRW Backgrounds

Electrodynamics

Maxwell's Theory:

$$S = -\frac{1}{4} \int d^4x F^{\mu\nu} F_{\mu\nu}, \quad F_{\mu\nu} = 2\partial_{[\mu} A_{\nu]}, \quad \text{invariant under } \delta A_\mu = \partial_\mu \Lambda.$$

Goal: Write Maxwell's theory in manifestly gauge invariant form.

Decomposition of vector field:

$$A_i = \hat{A}_i + \partial_i \psi, \quad \partial_i \hat{A}^i = 0, \\ \psi = \Delta^{-1}(\partial_i A^i), \quad \Delta \equiv \partial^i \partial_i \text{ invertible.}$$

Components of A_μ transform as: $\delta A_0 = \dot{\Lambda}$, $\delta \hat{A}_i = 0$, $\delta \psi = \Lambda$.

Maxwell's equations $\partial_\mu F^{\mu\nu} = 0$ in terms of \hat{A}_i and $\Phi \equiv A_0 - \dot{\psi}$:

$$\Delta \Phi = 0 \quad \rightarrow \quad \Phi = 0, \quad (\text{using } \Delta \text{ invertible}) \\ \square \hat{A}_i + \partial_i \dot{\Phi} = 0 \quad \rightarrow \quad \square \hat{A}_i = 0.$$

Manifestly gauge-invariant action: $S = \frac{1}{2} \int d^4x (\hat{A}^i \square \hat{A}_i - \Phi \Delta \Phi).$

Gravity on Flat Space

Einstein-Hilbert Action:

$$S = \int d^4x \sqrt{-g} R,$$

invariant under $\delta g_{\mu\nu} = \mathcal{L}_\xi g_{\mu\nu} \equiv \xi^\rho \partial_\rho g_{\mu\nu} + \partial_\mu \xi^\rho g_{\rho\nu} + \partial_\nu \xi^\rho g_{\mu\rho}$.
Expanding around flat space, $g_{\mu\nu}(t, x) = \eta_{\mu\nu} + h_{\mu\nu}(t, x)$:

$$S = -\frac{1}{2} \int d^4x h^{\mu\nu} G_{\mu\nu}(h), \quad \text{invariant under } \delta h_{\mu\nu} = \partial_\mu \xi_\nu + \partial_\nu \xi_\mu.$$

Decompose $h_{\mu\nu}$:

$$h_{00} = -2\phi,$$

$$h_{0i} = B_i + \partial_i B,$$

$$h_{ij} = \hat{h}_{ij} + \partial_i E_j + \partial_j E_i + \partial_i \partial_j E + \frac{1}{3} \delta_{ij} (C - \Delta E),$$

where $\partial^i B_i = \partial^i E_i = 0$, $\partial^i \hat{h}_{ij} = \delta^{ij} \hat{h}_{ij} = 0$.

Gravity on Flat Space

Decompose gauge parameter (ξ_0, ξ_i) : $\xi_i = \zeta_i + \partial_i \chi$ with $\partial^i \zeta_i = 0$.

$$\delta h_{00} = -2\delta\phi = 2\dot{\xi}_0,$$

$$\delta h_{0i} = \delta(B_i + \partial_i B) = \dot{\zeta}_i + \partial_i(\dot{\chi} + \xi_0),$$

$$\begin{aligned}\delta h_{ij} &= \delta[\widehat{h}_{ij} + \partial_i E_j + \partial_j E_i + \partial_i \partial_j E + \frac{1}{3}\delta_{ij}(C - \Delta E)] \\ &= \partial_i \zeta_j + \partial_j \zeta_i + 2\partial_i \partial_j \chi\end{aligned}$$

Gauge transformations of SVT components:

$$\delta\phi = -2\dot{\xi}_0,$$

$$\delta B_i = \dot{\zeta}_i, \quad \delta B = \dot{\chi} + \xi_0,$$

$$\delta E_i = \zeta_i, \quad \delta E = 2\chi, \quad \delta C = 2\Delta\chi, \quad \delta\widehat{h}_{ij} = 0.$$

Gauge invariant combinations (Bardeen variables):

$$\Sigma_i \equiv \dot{E}_i - B_i, \quad \Phi \equiv \frac{1}{3}(C - \Delta E), \quad \Psi \equiv \phi + \dot{B} - \frac{1}{2}\ddot{E}$$

Gravity on Flat Space

Einstein equations:

$$G_{00} = -\Delta\Phi,$$

$$G_{0i} = \frac{1}{2}\Delta\Sigma_i - \partial_i\dot{\Phi},$$

$$G_{ij} = -\frac{1}{2}\square\hat{h}_{ij} - \partial_i\partial_j\left(\Psi + \frac{1}{2}\Phi\right) + \delta_{ij}\left(\Delta\Psi + \frac{1}{2}\Delta\Phi - \ddot{\Phi}\right) + \partial_{(i}\dot{\Sigma}_{j)}.$$

Reduce to:

$$\square\hat{h}_{ij} = 0.$$

Rewrite Fierz-Pauli action in manifest gauge invariant form:

$$S = \int d^4x \left\{ \frac{1}{4}\hat{h}^{ij}\square\hat{h}_{ij} - \frac{1}{2}\Sigma^i\Delta\Sigma_i - \frac{1}{2}(4\Psi + \Phi)\Delta\Phi + \frac{3}{2}\Phi\ddot{\Phi} \right\}$$

Gravity on FRW Backgrounds

Einstein-Hilbert Action with Minimally Coupled Scalar Field:

$$S = \int d^4x \sqrt{-g} \left\{ R - \frac{1}{2} g^{\mu\nu} \partial_\mu \varphi \partial_\nu \varphi - V(\varphi) \right\},$$

Expand around FRW:

$$\begin{aligned} g_{\mu\nu}(\eta, x) &= a^2(\eta)(\eta_{\mu\nu} + h_{\mu\nu}(\eta, x)), \\ \varphi(\eta, x) &= \varphi_0(\eta) + \varphi_1(\eta, x). \end{aligned}$$

Gauge invariant variables:

$$\begin{aligned} \Sigma_i &= \dot{E}_i - B_i, \\ \Phi &= -C + \frac{1}{3} \Delta E - H(B - \dot{E}), \\ \Psi &= \phi + H(B - \dot{E}) + \dot{B} - \ddot{E}, \\ \Theta &= \varphi_1 + \dot{\varphi}_0(B - \dot{E}). \end{aligned}$$

Gravity on FRW Backgrounds

After lots of computation:

$$S = \int d^4x a^2 \left\{ \frac{1}{4} \dot{\hat{h}}^{ij} \dot{\hat{h}}_{ij} + \frac{1}{4} \hat{h}^{ij} \Delta \hat{h}_{ij} - \frac{1}{2} \Sigma_i \Delta \Sigma^i \right. \\ \left. + 4\Phi \Delta \Psi - 2\Phi \Delta \Phi - 6(\dot{\Phi} + H\Psi)^2 + \frac{1}{2} \dot{\varphi}_0^2 \Psi^2 \right. \\ \left. + \frac{1}{2} \dot{\Theta}^2 + \frac{1}{2} \Theta \Delta \Theta - \frac{1}{2} a^2 V'''(\varphi_0) \Theta^2 \right. \\ \left. + \dot{\varphi}_0 \Theta (\dot{\Psi} + 3\dot{\Phi}) - 2a^2 V'(\varphi_0) \Theta \Psi \right\}.$$

Questions:

- How to discern propagating degrees of freedom?
- How to expand to cubic order (and beyond!) in an efficient way?